

Pitfalls in Making Pretty Pictures

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High-quality images for public outreach pushed by

- STScI/STECF, Hubble Heritage Project
- Jean-Charles Cuillandre, CFHT
- Adam Block, NOAO
- ...

Large investment of time, long learning curve

If you find your own images here: I'm sorry!

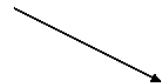
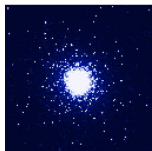
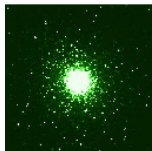
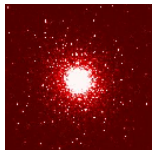
Content

- 1 Overview
 - Things that usually go wrong
 - Essentials for a good picture

- 2 Guidelines for pretty pics
 - Observations
 - Data reduction
 - Photoshop science

1. Overview

It looks all so simple...



Things that usually go wrong

In the order encountered
by beginners:

- **Dynamic range compression:**
Saturation
Background clipping
- Distortion correction
- Colour calibration
- Over-processing and under-exploitation



(C) ESO, M83, SOFI

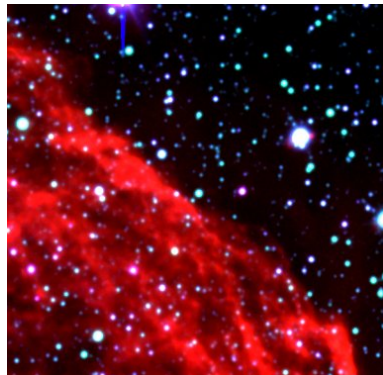
Use ESO **FitsLiberator**

(Lars Lindberg Christensen)

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(C) ESO, Part of LMC, WFI

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(C) ESO, RCW108, WFI

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(C) Jim Misti, M63

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Do it right from the beginning, and your PR officer will love you

Essentials for a good pretty picture

Observing

- A good observing strategy

Data reduction

- Get the relative astrometry right
- Careful sky background modelling for extended targets and multi-chip cameras
- **Absolute colour calibration**
(5% of all astronomers are colour blind)

Photoshop science

- Dynamic range compression
- Avoid over-processing

2. Guidelines – Observations

No good observing strategy:

No good science

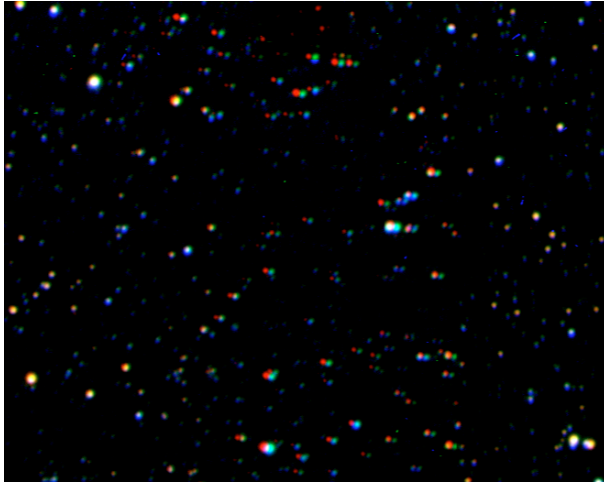
No nice pics

Please do:

- Extensive (excessive) dithering:
Defects, defringing, inter-chip gaps, astrometry,
photometry
- Blank fields for extended targets, not only in the near-IR:
Defringing, sky modelling

This holds in particular for multi-chip cameras.

2.1 Guidelines – Astrometry



(C) ESO, Capodimonte Deep Field, WFI

2.2 Guidelines – Sky subtraction

Colour pictures are excellent indicators of data reduction quality



M 83, near-IR

2.2 Guidelines – Sky subtraction

Colour pictures are excellent indicators of data reduction quality



M 83 (showing sub-percent residuals in sky background)

☐ Model the sky
☒ Subtract a constant sky

DT BMIN SIZE

☒ Smoothing with SEXtractor
☐ Smoothing with a Gaussian

Choose statistics
Median
From chip 3

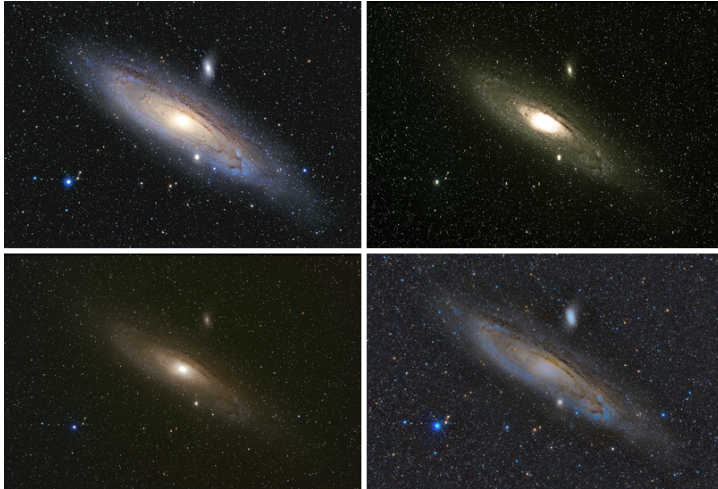
Region with empty sky (optional)
RA 1 xmin
RA 2 xmax
DEC 1 or ymin
DEC 2 ymax

Manual override (individual exposures)

Show mode
Use all chips
Show mosaics only
Image name Mode

THELI sky background configuration

2.3 Guidelines – Colour calibration



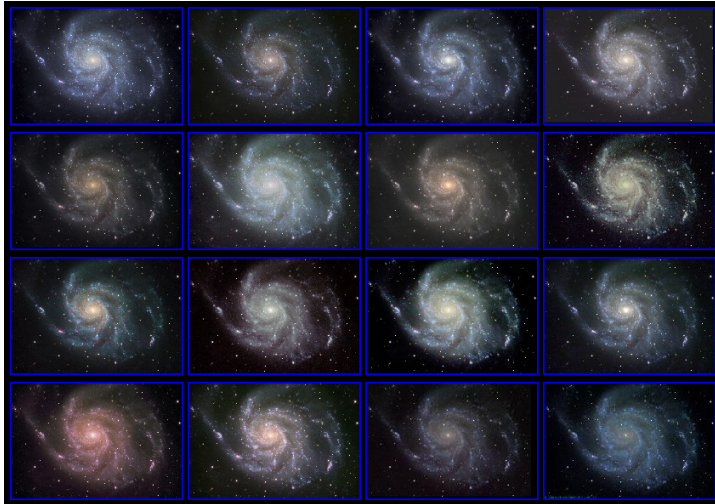
(C) M 31, Jim Misti

2.3 Guidelines – Colour calibration



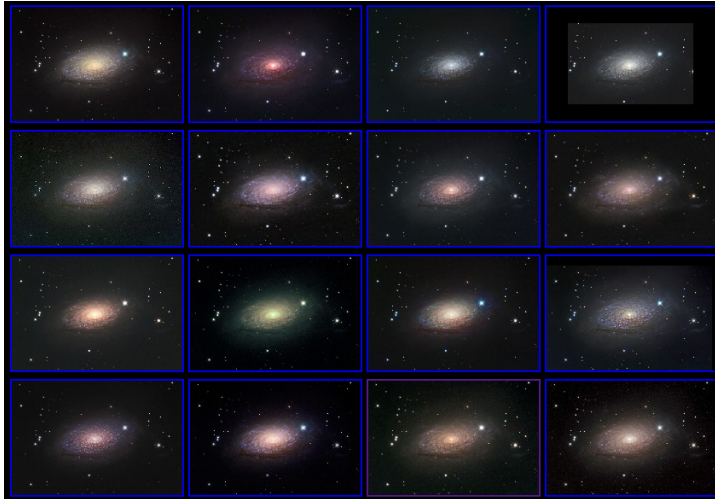
(C) Melotte 15, Jim Misti

2.3 Guidelines – Colour calibration



(C) M 101, Jim Misti

2.3 Guidelines – Colour calibration



(C) M 63, Jim Misti

2.3 Guidelines – Colour calibration

M 63 with SDSS calibration



2.3 Colour calibration – G2V stars

Requirement: an absolute white point

Observations of a G2V solar-type standard star.

But:

- G2 stars are rare
- If in field of view, then most likely saturated
- Nightly observations required to match atmosphere
- OPC/TAC and you won't like that waste of time
- Simply not available for archival data

2.3 Colour calibration – SDSS ugr calibration

Solution: Select G2V stars photometrically

SDSS

$$1.38 < u - g < 1.48$$

$$0.34 < g - r < 0.54$$

NOMAD

$$0.62 < B - V < 0.68$$

$$0.30 < V - R < 0.70$$

Advantages:

- Internal calibration
- Almost always in the field of view
- Comes for free, no extra time



2.3 Colour calibration – SDSS ugr calibration

The corresponding dialogue in THELI

Red image: red2_cropped.fits

Green image: green2_cropped.fits

Blue image: blue2_cropped.fits

Method: PHOTOMETRIC CATALOG

Photometric reference catalog: SDSS

u-g: 1.38 1.48

g-r: 0.34 0.54

G2-star defaults

	Factor	Error
Red	1.348	0.039
Blue	1.661	0.041
Green	1.000	0.000

Based upon 13 stars

Calibrate

Reset

Fall-back solution: all stars white on average

This works nicely in the near-IR :-)



(C) ESO, M83 spiral arm, VLT/FORS2



M83 spiral arm, VLT/HAWK-I (JHKs, $0.4''$ seeing)

(C) Yuri Beletsky, Mischa Schirmer, Mark Gieles



M83 core, HAWK-I

(C) Yuri Beletsky, Mischa Schirmer, Mark Gieles

2.4 Guidelines – Photoshop science

Typical post-processing tasks

- Dynamic range compression (e.g. through FitsLiberator)
- False-colour composition (e.g. more than 3 filters)
- Noise filtering
- Sharpening (wavelet, deconvolution)
- Background flattening

Last, but not least

- **Less is more, don't overdo it**
- **Give your eyes a rest**

Thank you